

Resonances for r -normally hyperbolic trapped sets

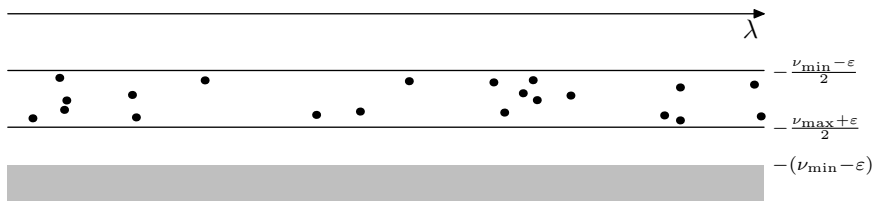
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June 17, 2013

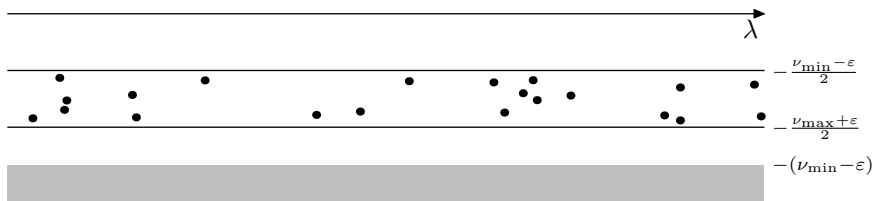
- **Resonances** are a discrete set of complex frequencies describing decay of waves/classical correlations for **open systems**
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- Our result relies on the construction of a microlocal projector corresponding to resonances in the band, a **Fourier integral operator**
- We also get new information on microlocalization of **resonant states**

Motivation: ringdown on black holes

Chandrasekhar '83:

... we may expect that any initial perturbation will, during its last stages, decay in a manner characteristic of the black hole itself and independent of the cause. In other words, we may expect that during these last stages, the black hole emits gravitational waves with frequencies and rates of damping that are characteristic of the black hole itself, in the manner of a bell sounding its last dying notes.

Motivation

Linear waves on (slowly) rotating Kerr–de Sitter black holes (\tilde{M}, \tilde{g})

$$\square_{\tilde{g}} u = f \in C_0^\infty, \quad u|_{t < 0} = 0.$$

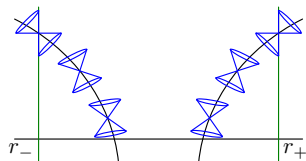
Open system: most energy escapes through the **event horizons**.

What is the behavior of $u(t)$ as $t \rightarrow +\infty$?

Theorem [D '10, '11]

There is a discrete set of **resonances** $\{\lambda_j\} \subset \{\text{Im } \lambda \leq 0\}$ such that $\forall \nu > 0$, the following **resonance expansion*** holds

$$u(t, x) = \sum_{\text{Im } \lambda_j > -\nu} e^{-it\lambda_j} \Pi_j f(x) + \mathcal{O}(e^{-\nu t}).$$



Two radial timelike geodesics, with light cones shown

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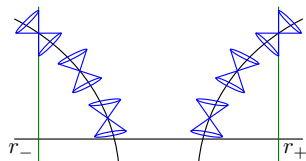
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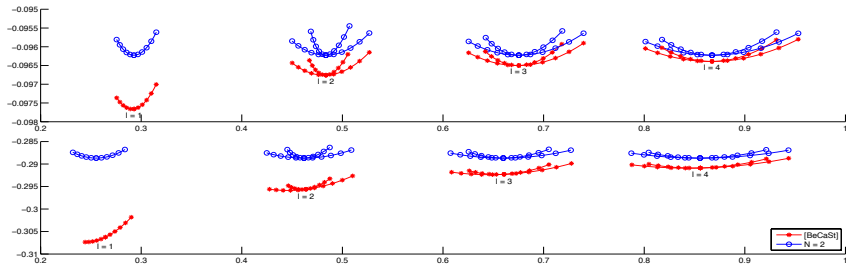


Two radial timelike geodesics, with light cones shown

Theorem [D '10, '11]

Moreover, resonances satisfy a **quantization condition**, that is they lie asymptotically on a distorted lattice.

The quantization condition agrees with the exact values computed in **Berti–Cardoso–Starinets '09** (shown here for $a = 0, 0.05, \dots, 0.25$)

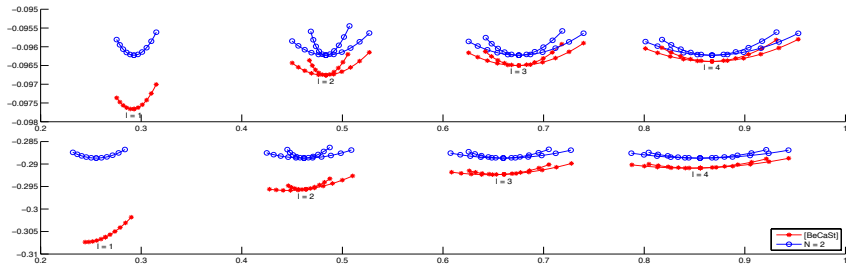


Previous work: Sá Barreto–Zworski '97, Bony–Häfner '07,
Dafermos–Rodnianski '07, Melrose–Sá Barreto–Vasy '08,
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How about (stationary) **perturbations** of Kerr–de Sitter?

Strategy for exact K–dS:



Perturbations destroy symmetries and our ability to handle individual resonances (similarly to the compact case).

Strategy for perturbations:



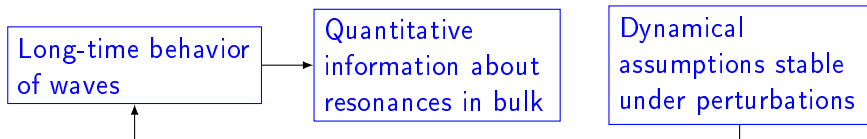
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Setup

The structure of open ends/event horizons does not change our result (Vasy '10 for event horizons of K-dS) \implies use a simpler model at infinity:

$$(\tilde{M}, \tilde{g}) = (M_x, g) \times \mathbb{R}_t, \quad (M_x, g) \simeq \mathbb{R}^n \text{ outside of a compact set}$$

Define resonances as poles of the meromorphic continuation of the resolvent of the Laplacian $\Delta \geq 0$:

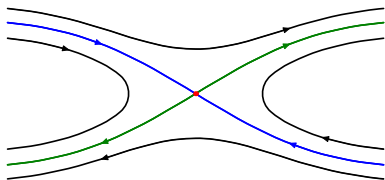
$$R(\lambda) = (\Delta - \lambda^2)^{-1} : \begin{cases} L^2(M) \rightarrow L^2(M), & \text{Im } \lambda > 0; \\ L^2_{\text{comp}}(M) \rightarrow L^2_{\text{loc}}(M), & \text{Im } \lambda \leq 0. \end{cases}$$

Long-living resonances: $\text{Re } \lambda \rightarrow \infty, |\text{Im } \lambda| \leq C$. Depend on trapping for the geodesic flow e^{tH_p} , $\rho(x, \xi) = |\xi|_g$:

$$\Gamma_{\pm} = \{\rho \in T^*M \mid e^{tH_p}(\rho) \not\rightarrow \infty \text{ as } t \rightarrow \mp\infty\}, \quad K = \Gamma_+ \cap \Gamma_-.$$

Dynamical assumptions

- Γ_{\pm} are C^r hypersurfaces ($r \gg 1$)
- $K = \Gamma_+ \pitchfork \Gamma_-$ is **symplectic**
- $\nu_{\min} \leq \nu_{\max}$ are the expansion rates of e^{tH_p} along Γ_{\pm} **transversally to K**
- $\mu_{\max} \geq 0$ is the maximal expansion rate of e^{tH_p} **along K**
- **r -normal hyperbolicity**: $\nu_{\min} > r \cdot \mu_{\max}$
- **pinching**: $\nu_{\min} > \frac{1}{2}\nu_{\max}$

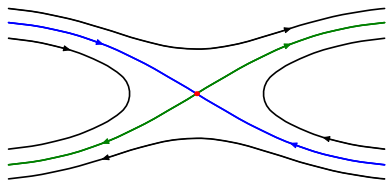


Normally hyperbolic trapping,
with Γ_+ , Γ_- , and K .

- Stable under small smooth perturbations: **Hirsch–Pugh–Shub '77**
- True for slowly rotating K–dS (for S–dS, K is the **photonsphere**)
- More general **normally hyperbolic** trapping ($r = 0$) appears for **Ruelle resonances for Anosov flows [Faure–Sjöstrand '11]** and **quantum chemistry [Goussev–Schubert–Waalkens–Wiggins '10]**

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Main theorem [D '13]

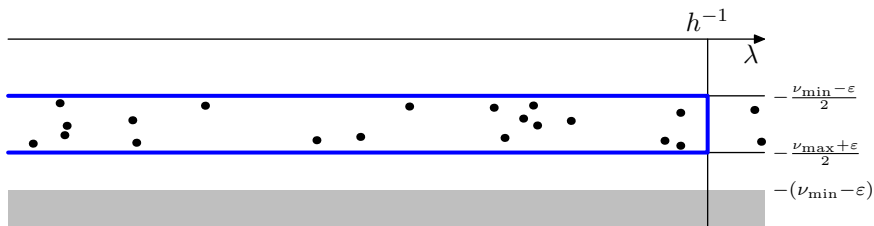
For each fixed $\varepsilon > 0$, there exist **two** resonance free strips

$$\{\operatorname{Re} \lambda \gg 1, \operatorname{Im} \lambda \in [-(\nu_{\min} - \varepsilon), -\frac{1}{2}(\nu_{\max} + \varepsilon)] \cup [-\frac{1}{2}(\nu_{\min} - \varepsilon), \infty)\}$$

and in these strips, $\|\chi R(\lambda) \chi\|_{L^2 \rightarrow L^2} = \mathcal{O}(1)$ for $\chi \in C_0^\infty(M)$.

Moreover, resonances between the two strips satisfy a **Weyl law** as $h \rightarrow 0$:

$$\#\{0 \leq \operatorname{Re} \lambda \leq h^{-1}, \operatorname{Im} \lambda > -\frac{1}{2}(\nu_{\max} + \varepsilon)\} = \frac{\operatorname{Vol}_\sigma(K \cap B^*M)}{(2\pi h)^{n-1}} + o(h^{1-n}).$$



What about exact Kerr–de Sitter?

- $M > 0$ mass of the black hole
- a speed of rotation, $|a| < M$
- $\Lambda > 0$ cosmological constant,
 $9\Lambda M^2 < 1$

The metric is r -normally hyperbolic
for $\Lambda \ll 1$ or $|a| \ll 1$: $\nu_{\min} > 0$,
 $\mu_{\max} = 0$.

However, the pinching condition $\nu_{\min} > \frac{1}{2}\nu_{\max}$ breaks down for $a = M - \varepsilon$,
 $\varepsilon \ll 1$: there exist two closed trajectories (equators) $E_{\pm} \subset K$ such that the
local expansion rate ν satisfies as $\varepsilon \rightarrow 0$

$$\nu \sim \frac{3\sqrt{3}}{28M} \quad \text{on } E_+, \quad \nu \sim \frac{\sqrt{\varepsilon/2M}}{M} \quad \text{on } E_-.$$

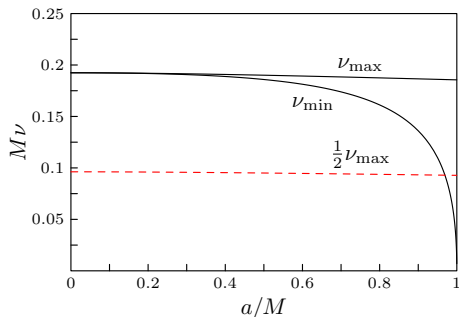
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Previous work

- Resonant free strips for NH trapped sets: Gérard–Sjöstrand '88, Dolgopyat '98, Liverani '04, Tsujii '10, Wunsch–Zworski '11, Nonnenmacher–Zworski '13. Our gap $\{\operatorname{Im} \lambda > -\frac{1}{2}(\nu_{\min} - \varepsilon)\}$ coincides with the one of N–Z '13 (valid in more general situations)
- Our Weyl law saturates general upper bounds in strips: Sjöstrand '90, Guillopé–Lin–Zworski '04, Sjöstrand–Zworski '07, Nonnenmacher–Sjöstrand–Zworski '11, '12, Datchev–D '12, Datchev–D–Zworski '12.
- Completely integrable cases: Gérard–Sjöstrand '87, Christianson '07, Sá Barreto–Zworski '97, D '11
- Bands of resonances/Weyl laws: Zworski '87, Sjöstrand–Vodev '97, Sjöstrand–Zworski '99, Sjöstrand '00, '11, Faure–Tsujii '12, '13

F–T '13 shares some ideas with presented work. It applies in settings with lower regularity, but does not recover microlocal structure away from K .

Ingredients of the proof

- Fourier integral operator Π projecting microlocally onto resonant states for the band of resonances
- Left/right ideals of pseudodifferential operators annihilating Π
- Local reduction to Taylor expansion
- Positive commutator estimates (replaced in this talk by wave propagation estimates)
- Grushin problems: resonances as zeros of a Fredholm determinant
- Trace formulas above and below the band of resonances, using microlocal analysis in λ
- Complex analysis (almost analytic continuation and argument principle) to obtain the Weyl law

Model case

$$M = \mathbb{R}, \quad P_0 = xD_x + \frac{1}{2i}, \quad U_0(t)f(x) = e^{-t/2}f(e^{-t}x)$$

$$p(x, \xi) = x\xi, \quad \nu_{\min} = \nu_{\max} = 1$$

The Taylor expansion

$$f(e^{-t}x) = f(0) + \mathcal{O}(e^{-t})_{L^2_{\text{loc}}}$$

can be viewed as a resonance expansion:

$$U_0(t)f = U_0(t)\Pi_0 f + \mathcal{O}(e^{-3t/2}),$$

$$\Pi_0 f(x) := f(0),$$

$$U_0(t)\Pi_0 f = e^{-t/2}\Pi_0 f$$

A resonance at $\lambda = -\frac{i}{2}$ and $\Pi_0 = \delta_0 \otimes \mathbf{1}$ is the projector onto the corresponding resonant state.

General case

$$P = \sqrt{\Delta}, \quad U(t) = e^{-itP}$$

“Projector”:

$$\Pi : C_0^\infty(M) \rightarrow C^\infty(M), \quad \text{Fourier integral operator}$$

“Resonance expansion”: for $f \in C_0^\infty(M)$ living at frequencies $\sim h^{-1}$ and $\chi \in C_0^\infty(M)$, $\chi = 1$ near K , modulo $\mathcal{O}(h^\infty)$ errors,

$$\begin{aligned} \|\chi U(t)(1 - \Pi)f\|_{L^2} &= \mathcal{O}(h^{-1} e^{-(\nu_{\min} - \varepsilon)t}) \|f\|_{L^2}, \\ e^{-\frac{\nu_{\max} + \varepsilon}{2}t} \|\chi \Pi f\|_{L^2} &\lesssim \|\chi U(t) \Pi f\|_{L^2} \lesssim e^{-\frac{\nu_{\min} - \varepsilon}{2}t} \|\chi \Pi f\|_{L^2} \end{aligned}$$

This is enough for **two resonance free strips**

+ $u = \Pi u + \mathcal{O}(h^\infty)$ for resonant states in the band.

Also true for Kerr black holes, using **Vasy–Zworski** '00 (note: high frequency régime!)

The projector Π

Construct unique h -Fourier integral operator Π , associated to a canonical relation $\Lambda \subset T^*M \times T^*M$, such that:

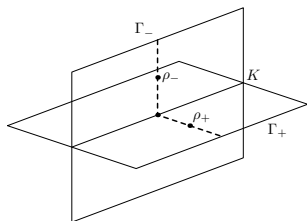
$$\Pi^2 = \Pi + \mathcal{O}(h^\infty), \quad (1)$$

$$[\Delta, \Pi] = \mathcal{O}(h^\infty), \quad (2)$$

microlocally near $K \cap S^*M$.

If $\mathcal{V}_\pm \subset T\Gamma_\pm$ are the symplectic complements of $T\Gamma_\pm$ in $T(T^*M)$ and $\pi_\pm : \Gamma_\pm \rightarrow K$ are the projections along the flow lines of \mathcal{V}_\pm , then

$$\Lambda = \{(\rho_-, \rho_+) \mid \rho_\pm \in \Gamma_\pm, \pi_-(\rho_-) = \pi_+(\rho_+)\}.$$



The canonical relation Λ (flow lines of \mathcal{V}_\pm dashed).

Transport equations

To construct Π , we need to solve **transport equations** of the form

$$H_p a = f, \quad a \in C^\infty(\Gamma_\pm), \quad a|_K = 0$$

where $f \in C^\infty(\Gamma_\pm)$, $f|_K = 0$.

The unique solution is given by the integral

$$a = \pm \int_0^\infty f \circ e^{\mp t H_p} dt.$$

The integral converges: $|f \circ e^{\mp t H_p}| = \mathcal{O}(e^{-(\nu_{\min} - \varepsilon)t})$, but is a smooth?

Differentiate along K :

$$\partial^k a = \pm \int_0^\infty \partial^k f \circ e^{\mp t H_p} \cdot \partial^k e^{\mp t H_p} dt.$$

Therefore, **r -normal hyperbolicity** guarantees that $a \in C^r$.

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$$P = \sqrt{\Delta}, \quad U(t) = e^{-itP}, \quad \Pi \in I(\Lambda);$$

$$U_0(t)f(x) = e^{-t/2}f(e^{-t}x), \quad \Pi_0 f(x) = f(0).$$

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$$\|\chi U(t)(1 - \Pi)f\|_{L^2} = \mathcal{O}(h^{-1}e^{-(\nu_{\min} - \varepsilon)t})\|f\|_{L^2}, \quad (3)$$

$$e^{-\frac{\nu_{\max} + \varepsilon}{2}t}\|\chi \Pi f\|_{L^2} \lesssim \|\chi U(t)\Pi f\|_{L^2} \lesssim e^{-\frac{\nu_{\min} - \varepsilon}{2}t}\|\chi \Pi f\|_{L^2}. \quad (4)$$

To prove (3) for the model case, write

$$(1 - \Pi_0)f = \chi f_1, \quad f_1(x) = \frac{f(x) - f(0)}{x}, \quad \|f_1\|_{L^2} \lesssim \|f\|_{H^1};$$

$$\chi U_0(t)(1 - \Pi_0)f = \chi U_0(t)\chi U_0(-t)U_0(t)f_1, \quad U_0(t)\chi U_0(-t) = e^{-t}\chi.$$

To prove (4) for the model case, note that $hD_x \Pi f = 0$ and thus $\langle a(x)\Pi f, \Pi f \rangle$ depends only on $\int a(x) dx$. Then

$$\|\chi U_0(t)\Pi f\|_{L^2}^2 = \langle \chi_t^2(x)\Pi f, \Pi f \rangle, \quad \chi_t = U_0(-t)\chi U_0(t),$$

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The general case is handled using the same ideas, but replacing x and hD_x by pseudodifferential operators Θ_- and Θ_+ solving

$$\Pi \Theta_- = \mathcal{O}(h^\infty), \quad \Theta_+ \Pi = \mathcal{O}(h^\infty). \quad (5)$$

Such Θ_\pm are not unique, but the sets of solutions to the equations form one-sided ideals invariant under the propagator $U(t)$.

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Open problems

- Show existence of further gaps and bands, below the first band of resonances. The main question is how to construct Fourier integral operators Π_1, Π_2, \dots projecting onto resonant states for each band.
- Obtain a quantization condition if the flow is completely integrable on the trapped set, recovering Gérard–Sjöstrand '87, Sá Barreto–Zworski '97, and D '11.
- Get an optimal remainder $\mathcal{O}(h^{2-n})$ in the Weyl law (we prove $o(h^{1-n})$ and $\mathcal{O}(h^{2-n-})$ should not be much harder), perhaps using the work of Sjöstrand '00 on the damped wave equation.

Thank you for your attention!